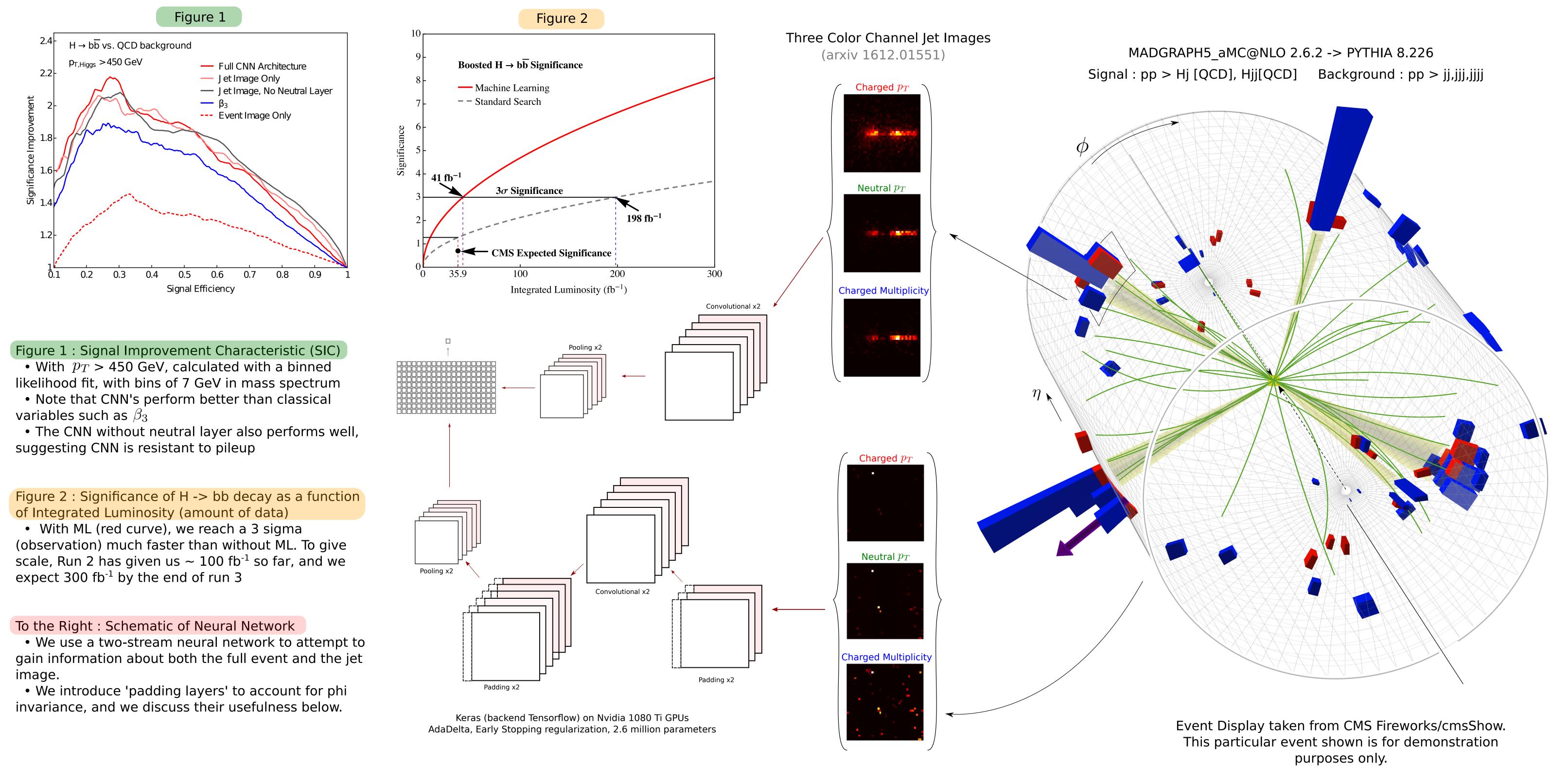
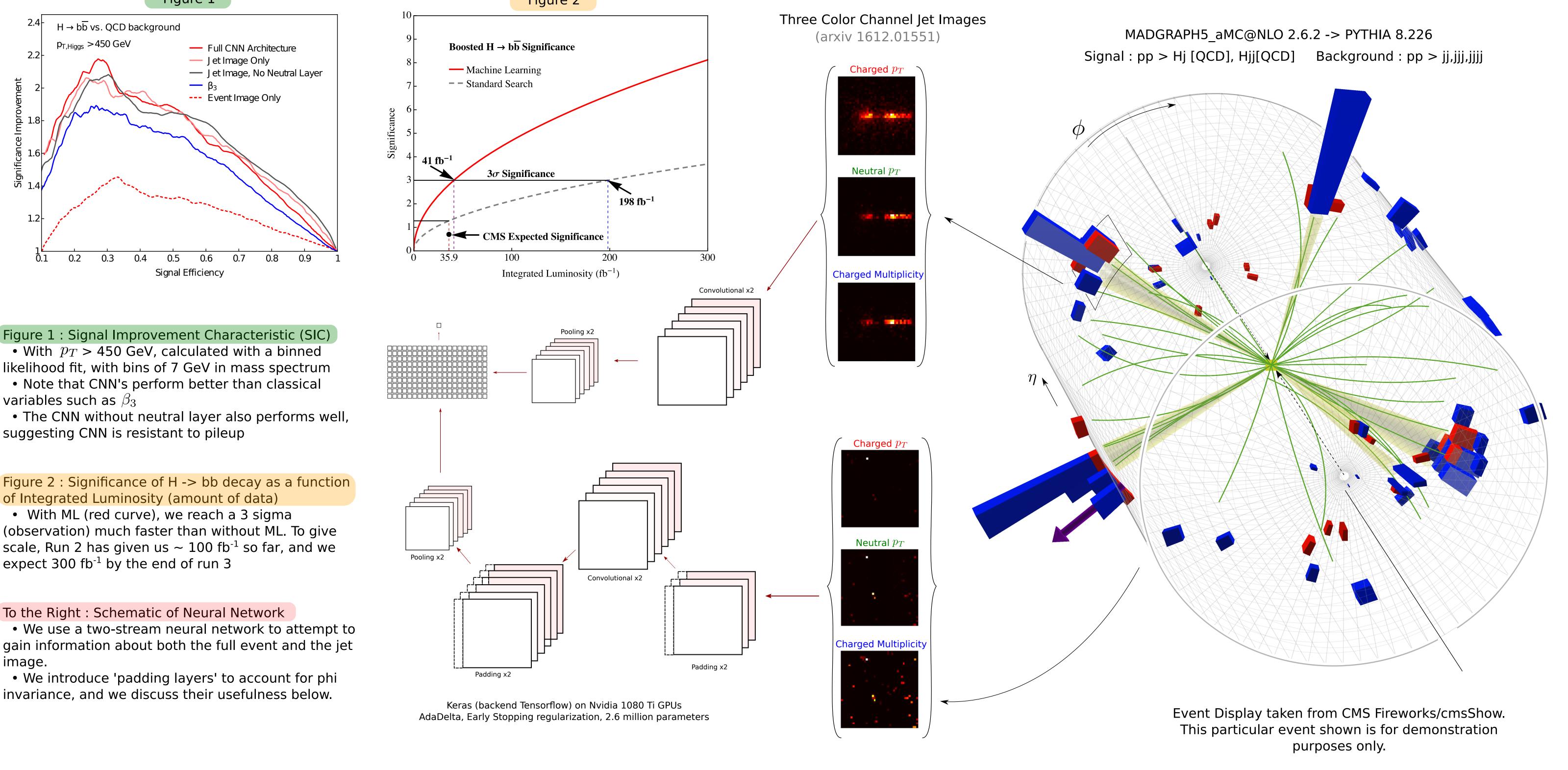
Boosting $H \rightarrow bb$ measurement with ML Joshua Lin^{1,2}, Marat Freytsis³, Ian Moult^{4,5}, and Benjamin Nachman¹

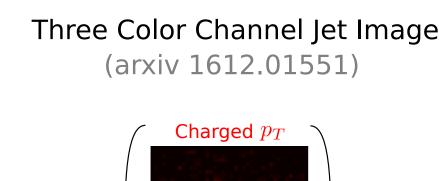


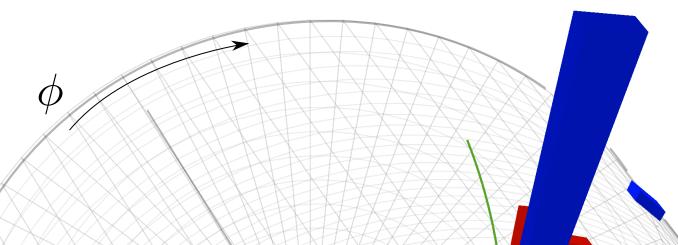
¹Physics Division, Lawrence Berkeley National Laboratory, Berkeley ²Department of Physics, University of California, Berkeley, Berkeley ³Institute of Theoretical Science, University of Oregon, Eugene ⁴Berkeley Center for Theoretical Physics, University of California, Berkeley ⁵Theoretical Physics Group, Lawrence Berkeley National Laboratory, Berkeley

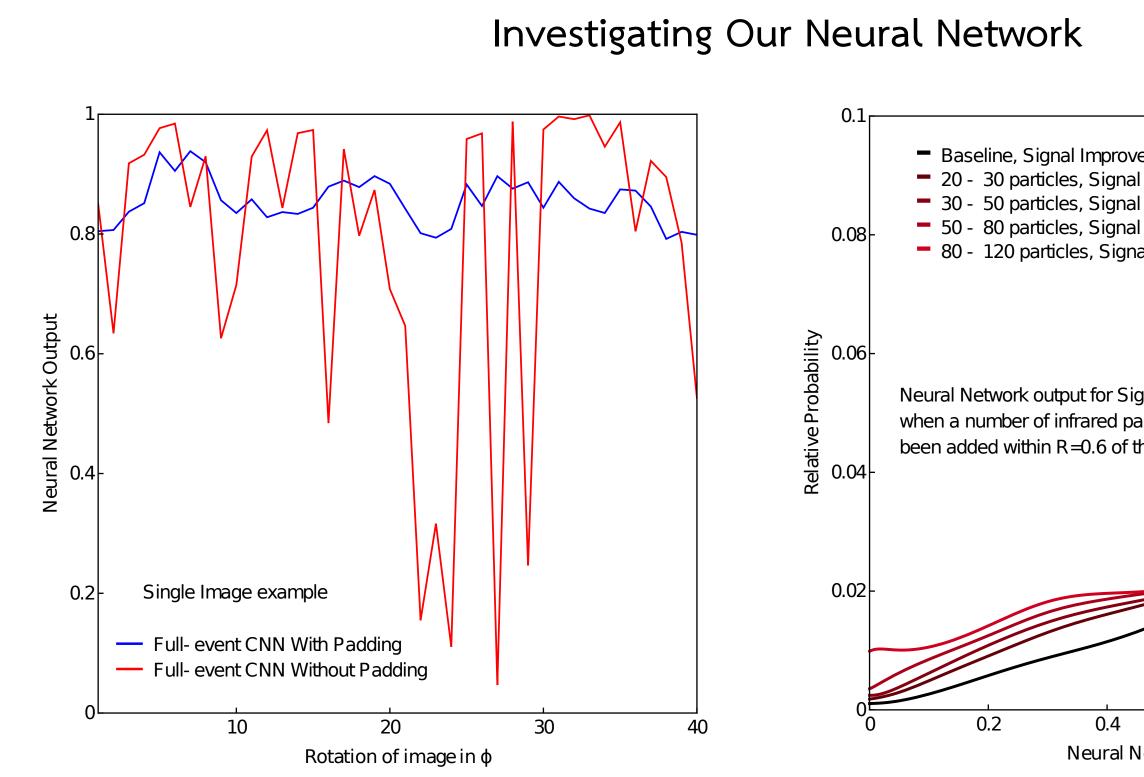












In this plot, we demonstrate that the padding layers on our Convolutional Neural Network have the desired effect of reducing the variance in the neural network output as the images are rotated in the phi direction.

The reason that the padded CNN is not completely rotationally invariant is that after all the convolutional/ pooling layers there is a dense layer that breaks the symmetry.

Bounds; using σ_{incl} and $\sigma_{p_T^{min} = 650 \text{GeV}}$ to break degeneracy $(3ab^{-1})$

0.0

 $Re(c_v) + c_H/2$

0.1

 $\sigma = 3$ with Machine Learning

 $\sigma = 3$ without Machine Learning

0.3 _T

0.2

0.1

0.0

-0.1

-0.2

-0.3 -

-0.4

-0.3

-0.2

-0.1

Baseline, Signal Improvement = 2.20 - 20 - 30 particles, Signal Improvement = 1.994 - 30 - 50 particles, Signal Improvement = 1.950 ■ 50 - 80 particles, Signal Improvement = 1.838 80 - 120 particles, Signal Improvement = 1.804 Neural Network output for Signal Images distribution, when a number of infrared particles (1 to 2 GeV) have been added within R=0.6 of the Higgs I et 0.8 0.6 Neural Network Output In this study, we tried adding different amounts of infrared noise to each of the pp collisions, by

generating some amount of 1 - 2 GeV pT particles and place them uniformly randomly in an 0.6×0.6 square around the Higgs Jet. We then feed these noisy images into the Neural Network that we had previously trained, and see how it affects the neural network performance. This plot shows that even at relatively high amounts of noise, we retain good signal improvement (> 1.8 factor)

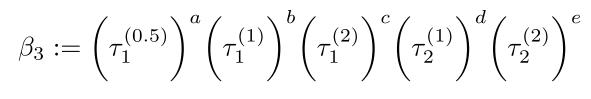
Classical Jet Substructure Variables

First, particles are usually clustered into jets, based on a specific choice of metric between particles. The algorithm proceeds by clustering the pair of particles closest to each other, until you remain with your jets. The most popular metric is shown right, with p = -1(anti-kT).

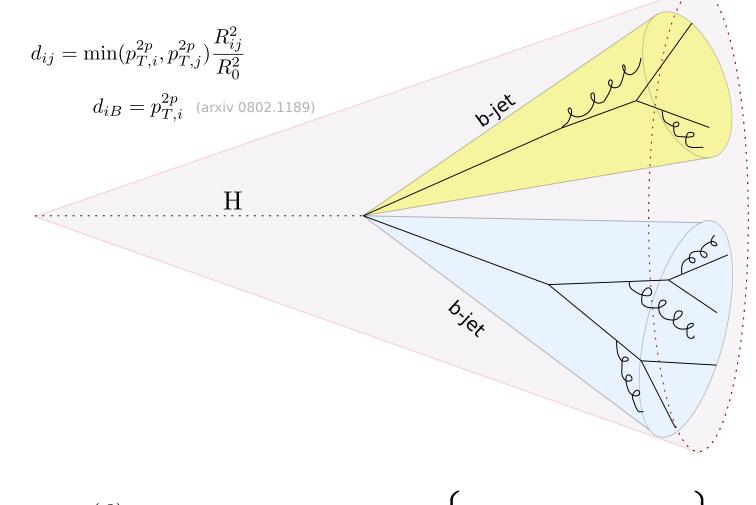
After clustering into jets, various physically motivated variables can be defined with input as a jet's constituent particles 4-vectors. These are mainly built with the intention of separating 'interesting jets' (Usually Higgs, vector bosons,..) from 'uninteresting jets' (QCD background)

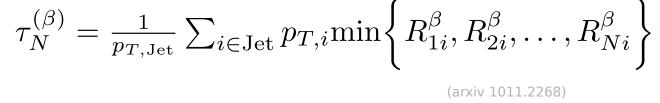
$$\vec{r}_i = (\Delta y_i, \Delta \phi_i) \quad \vec{t} = \sum_{i \in \text{Jet}} \frac{p_{T,i} |r_i|}{p_{T,\text{Jet}}} \vec{r}_i$$

The Jet Pull vector is a variable that attempts to capture the color flow information available in a jet. Specifically, it shows how much a jet (or subjet) is 'pulled off-center' by radiation; which is different for Higgs compared to QCD.

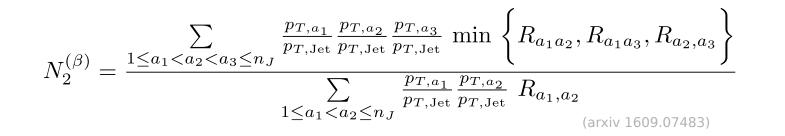


Built off the N-Subjettiness variables, the β_3 (modern) variable was built specifically to classify Higgs to bb jets apart from gluon to bb jets (which are the main source of QCD background in our study). The parameters are all optimized by means of a neural network to give optimal improvement in significance.

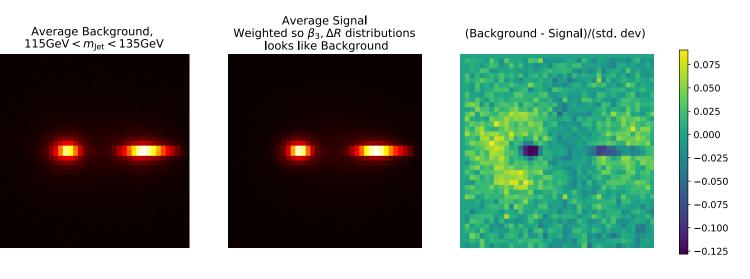




The N-subjettiness variables are designed to measure how much the jet looks like it is made out of N distinct components. They are useful in a wide variety of applications, specifically in this study as our signal involves a splitting of H to two b-quarks; we expect au_2 to be high.

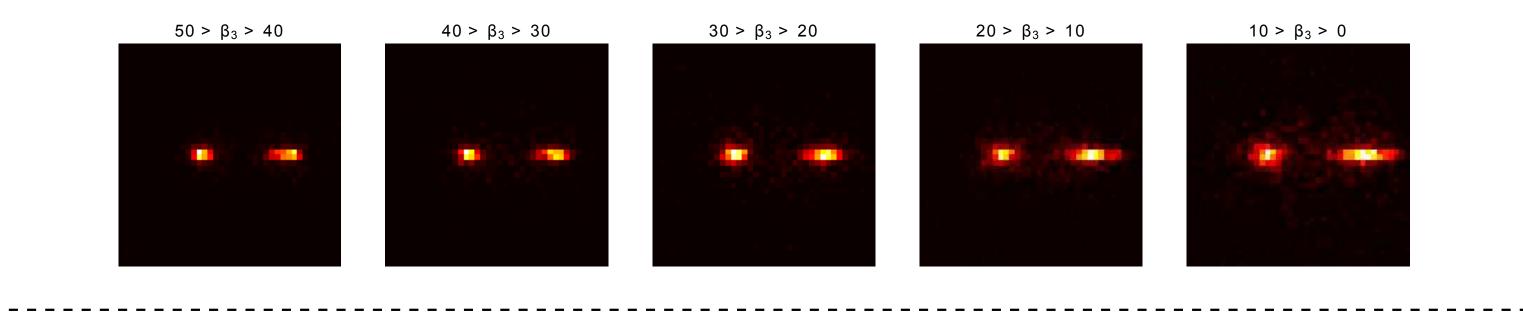


The N_2 variable is a (modern) discriminant built off of the energy correlation functions, which are designed to behave similarly to the N-subjettiness variables but with the benefit that they are defined without respect to a choice of subjet axes.



Left : To investigate what other information is available apart from β_3 , we can weight the signal images to have the same β_3 distribution as the background. After taking the difference, we find that there is still colour pull information to be learnt, shown by the radiation patterns remaining.

Bottom: Images to show what different values of β_3 correspond to.



0.2

0.3

Probing physics Beyond the Standard Model

 $\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} - \kappa_t \frac{m_t}{v} t\bar{t}h + i\tilde{\kappa}_t \frac{m_t}{v} \bar{t}\gamma_5 th + \kappa_g \frac{\alpha_s}{12\pi} \frac{h}{v} G^a_{\mu\nu} G^{\mu\nu a} + \tilde{\kappa}_g \frac{\alpha_s}{8\pi} \frac{h}{v} G^a_{\mu\nu} \tilde{G}^{a\mu\nu}$ $\kappa_t = 1 - \operatorname{Re}(c_y) - \frac{c_H}{2}, \quad \kappa_g = c_g, \quad \tilde{\kappa}_t = \operatorname{Im}(c_y), \quad \tilde{\kappa}_g = \tilde{c}_g$

Shown left are our bounds on the Wilson Coefficients that control the coupling of the Higgs Boson to tops and gluons. Changing these coefficients introduces two effects, one of which is an overall difference in the inclusive cross-section (which gives us the degeneracy band going from bottom-left to top right), and the other is an overall hardening of the p_T spectrum at high p_T , which allows us to 'break the degeneracy' and constrain the coefficients to a small region. We see that with ML, we achieve a much better constraint than without ML.